

## New Zealand's Warkworth Radio Astronomical Observatory

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#### Warkworth Radio Astronomy Observatory









3/80

AUT Radio Telescope

Auckland Ohena Ridge

Aotea Knoll

#### New Zealand's North Island

• Rotorua

North Island

☆Wellington









- Diameter: 12.1 m
- Manufacturer: Patriot/Cobham
- Shaped Cassegrain
- Slewing: 5 deg/s Az

1 deg/s El

- Surface: 0.35 mm (rms)
- Receivers/feeds
  - S/X (dual circular pols)
  - L (1.1-1.8 GHz; prime focus)
- H-maser (Symmetricom MH2010)
- Mk5B+ , Mk5C
- DBBC 2 (4 IFs)
- 10 Gbps network connectivity

6m L-band RT used in the first Trans-Tasman (New Zealand–Australia) VLBI observations in 2005.















### 



## Feed

- Coaxial S/X
  - S band 2.1 to 2.4 GHz
  - X band 8.1 to 9.1 GHz– RCP & LCP both bands
- ¼ wave plate polariser
  S Band
- Septum OMT polariser X Band







# L Band Feed

- Designed and manufactured by Intertronic
  Solutions Canada Peter Shields / Bill Imbriale
- Prime focus have to remove subreflector to use







### L Band Feed



### Receivers

- Room temperature un-cooled design
- SEFD
  - $\approx 3500 \text{ Jy} @ \text{S} \text{ Band}$
  - $\approx 3900 \text{ Jy} @ \text{X} \text{ Band}$



 Thanks to Peter McCulloch and UTAS for generous assistance with receiver development!





20/80

# **Rx System performance**

- System Equivalent Flux Density (SEFD) approx 3700 Jy
- Tsys approx 91 K
- Efficiency approx 0.65
- Pcal from Haystack

**Digital Backend DBBC2** 

- 4 x 512 MHz IFs
- 512 Mbps output in default modes
- 2 Gbps possible
- FILA10



# Digitiser + Recorders

- Data recorders:
  - Mk 5 B+ (2 Gbps to disk)
  - Mk 5 C (4 Gbps to disk)
  - Network streaming possible
    - Have done 512 Mbps eVLBI with data transport to Perth



#### 30 Metre



#### Specs:

Beam-waveguide cassegrain

Azimuth range: +/- 270 deg from East

Elevation range: 0 to 90 degrees

Slewing rate:0.3 deg/s (both El and Az)

Surface: 0.6 mm rms

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#### Conversion of a New Zealand 30-Metre Telecommunication Antenna into a Radio Telescope

Lewis Woodburn<sup>1</sup>, Tim Natusch<sup>1</sup>, Stuart Weston<sup>1,5</sup>, Peter Thomasson<sup>1,2</sup>, Mark Godwin<sup>3</sup>, Christophe Granet<sup>4</sup> and Sergei Gulyaev<sup>1</sup>

Description	Specification
System	Alt-azimuth, wheel-and-track, Cassegrain,
	beam-waveguide antenna
Drive system	Electric-servo, dual train for anti-backlash
Transmission frequency band	C-band
Reception frequency band	C-Band
Primary mirror diameter	30.48 m
Subreflector diameter	2.715 m
Azimuth maximum velocity in slew mode (open loop)	0.3 deg s <sup>-1</sup> or 18.0 deg min <sup>-1</sup>
Elevation maximum velocity in slew mode (open loop)	$0.3 \text{ deg s}^{-1} \text{ or } 18.0 \text{ deg min}^{-1}$
Max acceleration/deceleration in both axes	0.2 deg s <sup>-1</sup> s <sup>-1</sup>
Max tracking velocity (closed loop)	0.03 deg s <sup>-1</sup> or 1.8 deg min <sup>-1</sup>
	(estimated as no data in the NEC documentation)
Azimuth working range (as defined by soft limits)	-170° to 170°
Elevation working range (as defined by soft limits)	$0^{\circ}$ to $90^{\circ}$
Surface accuracy (rms)	0.4 mm
Track diameter	16.97 m
Total weight on track	268 tons
Wind speed in tracking operation	up to 40 m s <sup>-1</sup>
Survive wind speed in stow position	up to 70 m s <sup>-1</sup>

#### Table 1. Specifications of the Earth Station according to the manufacturer's (NEC) handbook.

Description	Specification
Azimuth maximum (tracking and slewing) velocity	0.37 deg s <sup>-1</sup> or 22.2 deg min <sup>-1</sup>
Elevation maximum (tracking and slewing) velocity	0.36 deg s <sup>-1</sup> or 21.6 deg min <sup>-1</sup>
Azimuth acceleration/deceleration	0.2 deg s <sup>-1</sup> s <sup>-1</sup>
Elevation acceleration/deceleration	0.25 deg s <sup>-1</sup> s <sup>-1</sup>
Azimuth working range (as defined by soft limits)	-179.0 to 354.0 deg
Elevation working range (as defined by soft limits)	6.0 to 90.1 deg

#### Table 3. New parameters after control system replacement.





















36/80
### 30 M optics + geometry



Figure 3-1. Cassegrain Antenna Geometry



#### Figure 1-1. Four-Reflector Primary Feed (YP-15412)





Figure 11. The C-band receiver with the new feedhorn transition unit manafactured by BAE Systems Australia. (Image courtesy of Stuart Weston.)





Figure 15. The 'First Light': the spectrum of the galactic Methanol Maser source G188.95+0.89 near 6.7 GHz.

Control system based on "Commercial Of The Shelf"

(COTS) industrial control technology Emerson Control Techniques













Zooming in to the heart of galaxy Centaurus A 14 million light-years away. Image Credits: Whole galaxy: I. Feain, T. Cornwell & R. Ekers (CSIRO/ATNF); ATCA northern middle lobe pointing courtesy R. Morganti (ASTRON); Parkes data courtesy N. Junkes (MPIfR). Inner radio lobes: NRAO / AUI / NSF. Core: S. Tingay (ICRAR) / ICRAR, CSIRO and AUT

## ASKAP & NZ VLBI of 1934-638

### Normal LBA at 1.4 GHz

## LBA with NZ and ASKAP





Image credit: Steven Tingay (see also Tzioumis et al. AJ, 140, 2010)



### Warkworth (WARK12M) is one of IVS network stations. And the only station in NZ that has more than one geodetic techniques



**PositioNZ** is LINZ's Global Positioning System Active Control Network. Through this site you of download GPS 30 second RINEX files from the active control stations which you can use with r GPS station data to determine precise positions in terms of New Zealand Geodetic Datum 2000

#### PositioNZ-RT – real-time data streams

Find out how to access real-time data from PositioNZ stations. more...

Text only version



Daily solutions for WARK





WARK is one of three NZ IGS stations, and the only one that operates all available GNSS systems

## GPS coordinates and Telescope survey

 Determination of VLBI reference point



- PositioNZ: WARK
- RTK GPS



- Local tie survey Dec 2012
- Recent survey Jan 2015

AuScope

## Study of Australian Tectonic Plate deformation



0/80

# **JAXA: IKAROS and Akatsuki**



# Akatsuki over NZ, August 2010



Resolution 1000 Hz, integration 500s

## Mars Express over NZ September 2010



Credit: Guifre Molera

53/80



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# **Precise mass determination and the nature of Phobos** by T. P. Andert et al.

We report independent results from two subgroups of the Mars Express Radio Science (MaRS) team who independently analyzed Mars Express (MEX) radio tracking data for the purpose of determining consistently the gravitational attraction of the moon Phobos on the MEX spacecraft, and hence the mass of Phobos. New values for the gravitational parameter  $(GM = 0.7127 \pm 0.0021 \times 10^{-3} \text{ km}^3/\text{s}^2)$  and density of Phobos (1876 ± 20 kg/m<sup>3</sup>) provide meaningful new constraints on the corresponding range of the body's porosity ( $30\% \pm 5\%$ ), provide a basis for improved interpretation of the internal structure. We conclude that the interior of Phobos likely contains large voids. When applied to various hypotheses bearing on the origin of Phobos, these results are inconsistent with the proposition that Phobos is a captured asteroid.

# **SpaceX Tracking**





57/80

# Ground tracking station

- Used SpaceX supplied TLE and files to generate ground track for Dragon passes
- Received downlink at 2.2 GHz
  - Flight and operational telemetry
  - Video (as seen on NASA TV)

- Track all passes visible from Warkworth
- Cover Launch, docking with ISS, opening of Dragon capsule
- Re-entry of Dragon
- SNR typically > 40 dB











# **Frequency Standard**

- Hydrogen Maser
  Symmetricom MMH2010
  - 5, 10, 100 MHz outputs
  - Allan deviation
    - 1s 2.0E-13
    - 1000s 3.2E-15
    - Floor 3.0E-15
  - Long term drift : <2.0E-16 per day</li>
- Temperature stabilised room: <+-0.5 deg.





## **Frequency Standard distribution**



#### Symmetricom Universal Time and Frequency Distributor

![](_page_65_Picture_1.jpeg)

## Maser reference distribution testing / performance test

![](_page_66_Figure_1.jpeg)

60.0

-90.0 105.0 120.0

150.0 165.0 180.0

67/80

### **REANNZ – Research and Education Network**

![](_page_67_Figure_1.jpeg)

- 10 Gbps backbone linking Universities and Crown research Institutes
- 40 Gbps International connectivity to Australia and US as of Oct 2015

![](_page_68_Figure_0.jpeg)

eVLBI, e-transfer via REANNZ has been demonstrated on several occasions.

Maximum data rates of 6.4 Gb/s where achieved from Warkworth for realtime e-transfer.

# **REANNZ** connectivity: timeline

- •2012: 10 Gbps
- •2014: 40 Gbps
- •2017: 80 Gbps
- •2022: 160 Gbps

# Noise floor measurements

![](_page_70_Figure_1.jpeg)

# **Collaboration with Japan**

• Establishment of the geodetic experiment environment

![](_page_71_Picture_2.jpeg)

 Experimental Phase\_cal thanks to NICT Kashima

![](_page_71_Picture_4.jpeg)

![](_page_71_Picture_5.jpeg)

Correlation processing Bandwidth synthesizing Data analysis

![](_page_71_Picture_7.jpeg)

![](_page_71_Picture_8.jpeg)

aseline length Ww-Ts 8,105 km Ww-K1 8,075 km Ww-K1 (2012/4/17) 8,075.003546 km
## Primary surface alignment

- Surface alignment conducted by photogrammetric testing; rms ≈ 0.35 mm
- Plan RF holography on surface in future to confirm the photogrammetry results and further refine if possible



0.04

0.03

0.02

0.01

-0.01

-0.02

-0.03

-0.04

## Integration of DBBC with Field System

- DBBC digitised channel power measurements fed back to Field System pointing model
- DBBC channel/IF configuration integration with Field System completed
- Thanks to Ed Himwich



74/80

## Astronomy Major



